

Deep learning methods for intra-day cloudiness prediction using geostationary satellite images in a solar forecasting framework

Franco Marchesoni-Acland^{a,d,*}, Andrés Herrera^b, Franco Mozo^b, Ignacio Camiruaga^c, Alberto Castro^{b,c}, Rodrigo Alonso-Suárez^{a,*}

^aLaboratorio de Energía Solar, Universidad de la República, Av. L. Batlle Berres km 508, Salto, Uruguay

^bInstituto de Ingeniería Eléctrica, Facultad de Ingeniería, UDELAR, J. H. y Reissig 565, Montevideo, Uruguay

^cInstituto de Computación, Facultad de Ingeniería, UDELAR, J. H. y Reissig 565, Montevideo, Uruguay

^dUniversité Paris-Saclay, ENS Paris-Saclay, CNRS, Centre Borelli, 91190, Gif-sur-Yvette, France

Abstract

Accurate solar resource forecasting remains a challenge. Electricity grid applications require both days-ahead and intra-day prediction. Satellite-based methods are known to be the best option for hourly intra-day solar forecasts up to some hours ahead. An adapted Deep Learning (DL) method has been recently reported to outperform the traditional Cloud Motion Vectors (CMV) strategy. This article analyzes the utilization of a well-documented computer vision DL architecture, the U-Net in various forms, for the satellite Earth albedo forecast problem (cloudiness), a straightforward proxy for solar irradiance forecast. It is shown that the U-Net performs better than advanced and optimized CMV techniques and previous art IrradianceNet, setting it at the state-of-the-art. The tests are done over the Pampa Húmeda region of southeast South America, an area in which challenging cloud conditions are frequent. The data for this study are GOES-16 visible channel images. These images present a finer spatial ($\simeq 1$ km/pixel) and temporal (10 minutes) resolution than previously explored data sources for solar forecasting. Moreover, the image size used here is $\times 4$ bigger (1024×1024 pixels) and the predictions reach further into the future (5 hours) than in previous works. The analysis includes several ablation studies, involving different architectures, optimization objectives, inputs, and network sizes. The U-Net is optimized for direct and differential image prediction, being the latter a better-performing option. More notably, the U-Net models are shown to be able to predict cloud extinction, something that has been a barrier for CMV methods.

Keywords: Solar forecast, U-Net, deep learning, satellite images, GOES-16 satellite.

1. Introduction

Solar photovoltaic (PV) is the main technology to convert the renewable power provided by the Sun into electricity (REN21, 2021). Due to the systematic cost reduction of this technology, its share in the world's

*Corresp. authors: R. Alonso-Suarez and F. Marchesoni-Acland, {r.alonso.suarez, marchesoniacland}@gmail.com

30 energy mix is continuously growing (IRENA, 2021). Regardless of its availability, high technology maturity
31 level, and low prices, solar energy is still a small portion of the world’s energy mix, partly due to the low
32 predictability of the solar resource (Voyant et al., 2017). The optimal operation of an electricity grid that
33 includes weather-dependent energy sources, such as solar PV and wind power, requires their forecast in
34 several time horizons. Solar forecasting in particular is mainly dependent on cloudiness forecasting, as the
35 other atmosphere variable components, for instance, water vapor and aerosol content, have a significantly
36 lower variability. The other source of solar variability is related to its well-known intra-day and seasonal
37 geometrical trend, governed by the Sun’s apparent movement, which can be effectively removed or modeled
38 (Lauret et al., 2022) by using a clear sky model (Lefèvre et al., 2013; Rigollier et al., 2004), leaving cloudiness
39 as the main driver of uncertainty in solar energy prediction.

40 Clouds undergo various transformations, such as formation, deformation, displacement, and dissipation,
41 which make predicting cloud cover a challenging task. Numerical Weather Prediction (NWP) models have
42 been historically used to provide solar irradiance forecasts by using their cloudiness prediction at different
43 levels in the atmosphere. These models numerically solve the equations that govern the atmosphere dynamic
44 and still are the only option to provide quality day(s)-ahead solar forecasts, typically augmented by post-
45 processing techniques (Yang et al., 2022). However, for solar irradiance intra-day forecasts, it has been shown
46 that pure satellite-based methods can achieve a better performance than regular NWP runs (i.e. not hourly
47 updated) for horizons up to 4-5 hours ahead (Kühnert et al., 2013; Perez et al., 2010). These satellite-based
48 methods have been mainly based on the two-dimensional cloud motion field estimation (Lorenz et al., 2004;
49 Peng et al., 2013; Cros et al., 2014; Urbich et al., 2019) known as Cloud Motion Vectors (CMV), from which
50 the clouds’ motion is extrapolated into the future. A detailed comparison between the main variants of
51 the CMV approach is presented in Aicardi et al. (2022), including the baseline block-matching technique of
52 Lorenz et al. and different types of optical flow methods (Horn and Schunck, 1981; Lucas and Kanade, 1981;
53 Farneäck, 2003; Zach et al., 2007) that are common in the computer vision field. Further, Aicardi et al.
54 bridged the gap between cloudiness and solar irradiance forecast performance evaluation, at least for the
55 inspected region (which is the same as the present work), showing that the best satellite-based forecasting
56 models at the cloudiness level are also typically the best at the solar irradiance level. The cloudiness level
57 in this context refers to the Earth’s albedo visible channel images (planetary reflectance), where clouds
58 are easily distinguishable. The albedo information is the common input for satellite-based solar irradiance
59 estimation methods (Perez et al., 2002; Rigollier et al., 2004; Qu et al., 2017; Laguarda et al., 2020). These
60 methods use updated formulations of the well-known satellite cloud index (Cano et al., 1986), calculated,
61 mainly, from visible channel images.

62 The CMV strategy presents some limitations at the current state of the art, such as the inherent two-
63 dimensional modeling of a three-dimensional problem, the motion extrapolation based on a static motion
64 field, and the inability to accurately forecast clouds’ formation and extinction, among others. For instance,

65 pure CMV strategies can only reallocate pixels of a satellite image but they are unable to create new values.
66 While this area advances, for instance through better ways to physically model the cloudiness evolution from
67 satellite images, data-intensive approaches have arisen in the field (Berthomier et al., 2020; Su et al., 2020;
68 Nielsen et al., 2021b). This is not surprising: deep learning (DL) methods have been recently introduced in
69 many meteorology-related areas (Ren et al., 2021; Ravuri et al., 2021; Espeholt et al., 2021; Schneider et al.,
70 2022) and DL methods excel when abundant pairs of inputs and targets are available, especially in the case
71 of automatically annotated data. The forecasting problem is inherently self-supervised, i.e. the targets are
72 naturally contained in the historical records and no manual annotations are needed. Visible channel satellite
73 images also provide large amounts of data, typically of dozens of TBs, from which the daylight cloudiness
74 behavior can be observed, modeled, and learned at a high rate (current geostationary meteorological satellite
75 scans the same Earth disk at a regular rate of 10-15 minutes). Due to all these, intra-day satellite cloudiness
76 forecasting is especially suited for deep learning methods.

77 1.1. Related Work

78 With the advent of deep learning, some research groups have started its utilization for black-box at-
79 mospheric modeling in problems related to cloudiness and solar irradiance forecasting. This subsection
80 describes the state-of-the-art work.

81 The U-Net architecture (Ronneberger et al., 2015) was used by Berthomier et al. (2020) for intra-day
82 cloud cover binary forecast over France up to 90 minutes ahead with a 15-minute time step. Berthomier et al.
83 showed that this type of U-Net cloud forecast is better than the forecast provided by the AROME NWP
84 model (Seity et al., 2011; Brousseau et al., 2016). Furthermore, it also compared favorably to different types
85 of neural networks, such as Recurrent Neural Networks (RNNs) and other Convolutional Neural Networks
86 (CNNs). The algorithms' training and validation were conducted for the binary cloudiness forecasting
87 problem (cloud vs. no cloud) using the Meteosat Second Generation (MSG) cloud classification product
88 as the input/output data, which is not directly translatable into a solar irradiance forecast. The satellite
89 images containing France's territory were cropped into 256×256 px patches that were processed separately
90 due to computational restrictions. Separated training and testing periods were considered, using one year
91 and a half for the former and six months for the latter. Even though the work did not aim to be used
92 for solar irradiance forecasting, as the use of binary cloud cover is limited for this task, it represented the
93 first deep learning attempt to cloud forecasting. In recent work, Nielsen et al. (2021b) proposed an adapted
94 deep learning architecture for solar irradiance forecast based on the MSG satellite information. In this case,
95 the European SARA-2.1 data set (Uwe et al., 2019) of satellite-derived global solar horizontal irradiance
96 was used. The architecture is similar to the one proposed by Tan et al. (2018) and was trained and tested
97 for a 4-hours ahead forecast with hourly time step. The deep learning algorithm outperformed the optical
98 flow TVL1 method (Zach et al., 2007; Sánchez et al., 2013) with the same satellite information, which has

99 been reported to be one of the best optical flow strategies for cloud and solar irradiance forecasting (Urbich
100 et al., 2019; Aicardi et al., 2022). The satellite images had European coverage with a 512×512 px resolution
101 and a 30-minute refresh rate. As in the work of Berthomier et al., the images were cropped, this time
102 into 128×128 px patches, and processed separately to avoid a high computational cost. The forecasts
103 were validated against the solar irradiance derived from satellite images and against four solar ground
104 measurement sites of the Baseline Solar Radiation Network (BSRN, <https://bsrn.awi.de/>) in Europe. An
105 important period was considered, using 5 years of images for training, 1 year for validation, and 2 years for
106 testing. The final results correspond to the evaluation across the latter split.

107 These works suggest that deep learning is currently the state-of-the-art best option for satellite intra-day
108 cloud cover and solar forecast. Furthermore, the two most performing architectures are the U-Net and
109 the IrradianceNet, which compare favorably to other architectures (Tan et al., 2018; Su et al., 2020). It is
110 not clear from the literature what is the best model for satellite Earth albedo and solar forecast, but it is
111 clear that they have outperformed NWP and CMV methods in each analyzed task and for the European
112 region. At the same time, well-performing optical flow CMV methods were inspected in Aicardi et al. (2022),
113 showing that a simple horizon-dependent spatial blurring of the CMV prediction can enhance performance.
114 As this performance booster was not considered in the previous deep learning vs CMV comparisons, it is
115 contemplated in the present work as an additional contribution. The present work benchmarks these three
116 techniques (U-Net, IrradianceNet, blurred CMV) for a region in South America with several ablation studies
117 and a common forecasting objective, showing that the U-Net is an upgrade of the current intra-day solar
118 forecasting performance. The U-Net is a performing and still widely used computer vision architecture
119 (Baranchuk et al., 2022; Rombach et al., 2022; Croitoru et al., 2023). It uses GOES-16 albedo images as a
120 regression target, as opposed to a classification target obtained by quantization (e.g. binarization). Finally,
121 it provides a novel scheme for the U-Net utilization in this framework (U-Net Diff), which resulted in the
122 best-performing strategy from the ones considered in this work, although closely followed by the traditional
123 U-Net.

124 In summary, IrradianceNet is the current performance lead on satellite-based solar forecasting using
125 deep learning, in fact being the only work providing in-depth evaluations. The present work addresses
126 current limitations, namely (i) moderate training and inference speed, (ii) limited image size, (iii) improvable
127 benchmarking, and (iv) laborious end-user implementation. These limitations are overcome by the U-Net
128 which is faster, extremely well-documented, can handle larger images (Figure 6), and is compared in this work
129 against blurred CMV, a tougher benchmark than plain CMV strategies (Figure 3). On top of overcoming
130 these limitations, the U-Net outperforms IrradianceNet in this region, setting it at the state-of-the-art and
131 providing a framework for future implementations of deep learning methods for this task.

132 *1.2. Article's outline*

133 This work builds upon previous works and compares past and concurrent research efforts in the field,
134 with a common ground in solar forecasting. It considers deep learning methods previously proposed in
135 different contexts and one of the best-performing CVM methods, including the important spatial blurring
136 step. The analysis is done with one complete year of satellite images over the Pampa Húmeda region of
137 South America, an area with challenging clouds' behavior and intermediate solar irradiance variability. The
138 results confirm that, for this region, deep learning methods are currently preferable over CMV methods.
139 Additionally, it is shown that the simple and lightweight U-Net architecture yields a strong performance
140 that surpasses all other methods. Regardless of its good performance, the main advantage remains practical:
141 the U-Net is a battle-tested architecture featured in many tutorials and articles in diverse research areas,
142 hence it is simpler to implement, train and deploy than custom architectures. Notwithstanding, there are
143 many details about the data processing, the training procedure, and the particular solar forecasting topic
144 that still need to be considered. Knowing which well-documented architecture performs well and how to
145 adapt it to the specific forecasting problem, makes deep learning solutions much more accessible to other
146 research teams and regions. The contributions of this work are as follows:

- 147 1. Illustrating the successful use of deep learning methods for 5-hours ahead albedo forecasting using
148 GOES-16 satellite images, being this the first comprehensive work aimed at solar satellite-based fore-
149 casting using this new satellite and over a region other than Europe. The only exception, to the best
150 of our knowledge, is the preliminary work of [Alonso-Suárez et al. \(2021\)](#), which presents the evaluation
151 of the operational CMV solar forecasting system developed and operated by our R&D group using
152 GOES-16 satellite images.
- 153 2. Showing that the U-Net deep learning model achieves the best performance for the albedo forecasting
154 problem when compared to Persistence, CMV, blurred CMV, and IrradianceNet. This benchmarks
155 all these models against the same data set over the same region, periods, and forecast horizons. The
156 findings demonstrate that the simpler U-Net architecture achieves the highest reported performance
157 in the field. This is also the first application of this algorithm in the specific framework of cloud albedo
158 prediction for solar irradiance forecasting.
- 159 3. Providing numerous ablation experiments for the specific problem and the U-Net strategy, regarding
160 the inference modality, the objective function, data augmentation, inner architecture, network size,
161 and alternative inputs. One of the tested U-Net configurations resulted in the best option.
- 162 4. Training and applying the previously proposed IrradianceNet algorithm over South America, thus
163 corroborating that it improves over the CMV and showing that it can handle different geographical
164 regions and image types with solid performance.
- 165 5. Displaying qualitative examples illustrating deep learning models being capable of predicting complex

166 phenomena such as cloud extinction, which regular CMV strategies are unable to perform.

167 The rest of this article is organized as follows. [Section 2](#) presents the satellite data set, including its
168 filtering and pre-processing. The metrics required for the work are introduced in [Section 3](#), and the considered
169 models along with their experiments are described in [Sections 4](#) and [5](#), respectively. This includes training
170 details and ablation studies. [Section 6](#) presents the final results over the test set, which is left aside from
171 the previous analysis. It also discusses some case studies in which the U-Net deep learning method was able
172 to forecast cloud extinction. Finally, [Section 7](#) summarizes the main conclusions of the work and provides
173 some ideas on how to further advance in this field, based on our experience and the observed limitations of
174 the techniques.

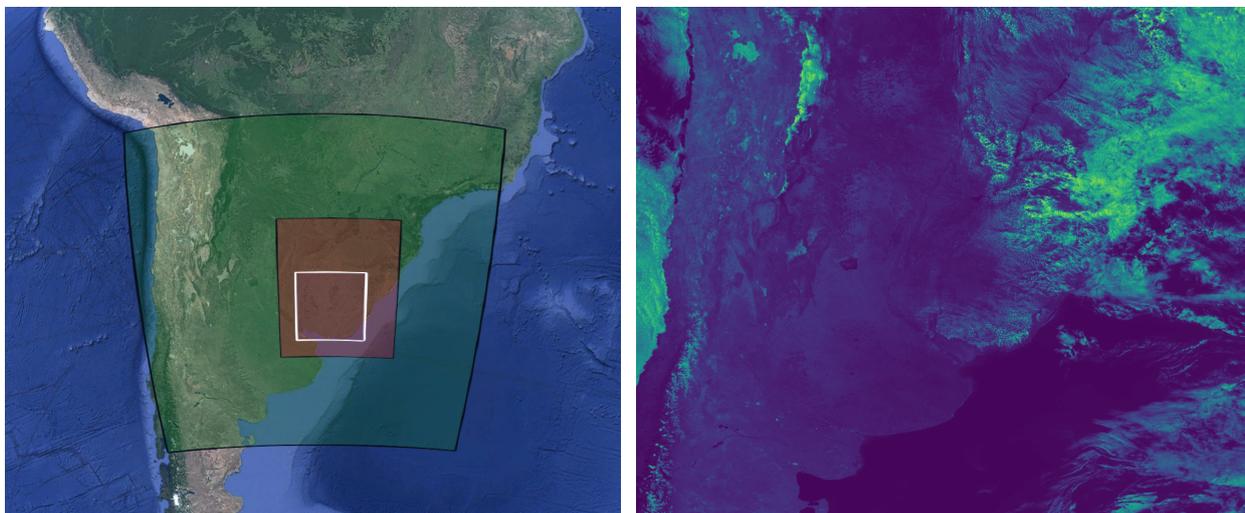
175 2. Data

176 2.1. Satellite images

177 The data set used in this work is managed and stored by the Solar Energy Laboratory (LES) of the
178 Universidad de la República (Udelar), Uruguay. It is composed of visible channel images taken from the
179 GOES-16 satellite’s C02 channel, which has its central wavelength at 640 nm. This is a meteorological
180 satellite administrated by the US’s National Oceanic and Atmospheric Administration (NOAA). It is located
181 in the geostationary position known as GOES-East (-75°W over the Earth’s equator) and its images are
182 freely available via different download mechanisms. This satellite provides a higher frame rate and resolution
183 for South America than the previous generation, with the important added value of schedule regularity. The
184 images are available every 10 minutes with a nominal resolution of 500 m at the satellite’s nadir. The visible
185 channel is useful for cloud detection as they typically reflect more sunlight than the ground, appearing
186 brighter in the images. This assumption does not hold for high albedo terrains, i.e. areas containing snow
187 or salt flats, which are not present in the portion of the images used in this work.

188 A large orthorectified crop of the satellite full disk images is processed by the LES lab for solar resource
189 assessment and forecasting, covering the southeast of South America as shown in the green rectangle of
190 [Figure 1a](#). An example of a satellite image from this database is shown in [Figure 1b](#), being of 2501×3001 px
191 and ranging 25 degrees in latitude and 30 degrees in longitude. However, due to computational restrictions
192 of the deep learning training, a reduced crop of 1024×1024 px is used in this work, shown in [Figure 1a](#)
193 as a red rectangle with a black frame. Note that this region contains at least $4\times$ more pixels than any
194 previous work involving DL in this context. All methods will be evaluated inside this region, thus having
195 the same space input information. This implies a built-in difference concerning the borders’ effect. In some
196 borders, and especially for the larger time horizons, the CMV is unable to provide predictions, as this would
197 require pixels from outside the image. Such regions are not considered for the CMV evaluation. The option
198 of providing a bigger frame for CMV forecast would result in this method having more spatial information

199 than the others. On the other hand, deep learning methods, by construction, provide a forecast for the whole
 200 image including these borders. This requires deep learning methods to figure out the most likely forecast
 201 based on the previous history of those pixels in similar situations. In this way, each method is evaluated in
 202 the pixels that it can predict and the space information given to all the methods is the same. Figure 1a also
 203 shows a smaller white-framed crop inside the red rectangle that will be used only for some comprehensive
 204 experiments on the U-Net architecture and related implementation details (Section 5). The size of this crop
 205 was 512×512 pixels, representing a region of 5×5 degrees in latitude and longitude. These crops are referred
 206 to as *smaller-crop data set*, which should not be assumed unless specified in the text. The images have a
 207 regular uniformly-spaced grid in the latitude-longitude domain with an average pixel side of $\simeq 1$ km in all
 208 these regions. The work was conducted with 10-minute images taken from the whole year 2020.



(a) Location of the satellite crops.

(b) Example image of the LES satellite database.

Figure 1: Satellite area considered in the work. It represents the southeast of South America, known as Pampa Húmeda.

209 Complementary to the satellite image data, an elevation map of the region is used for some experiments,
 210 as it is recommended by Nielsen et al. for the IrradianceNet method. Note that the southeast side of the
 211 South American continent is mostly flat. This region is referred to as Pampa Húmeda and it is considered
 212 warm, temperate, and humid, with hot summers, being Cfa in the updated Köppen-Geiger climate classifi-
 213 cation (Peel et al., 2007). The area is characterized by challenging mesoscale convective systems that tend
 214 to peak during daytime (Salio et al., 2007; Rasmussen et al., 2014) and has no rainy or dry season, being
 215 the cloudiness distributed throughout the year. In terms of ground-level solar irradiance, the short-term in-
 216 termittency is intermediate (Alonso-Suárez et al., 2020), meaning that clear-sky, partly cloudy, and overcast
 217 conditions are all present and frequent.

218 2.2. Filtering and preprocessing

219 The original data set comprises 52314 images (daytime and nighttime). These cover almost all the 10-
220 minute intervals over the year (only 390 images are missing). The original data set was filtered in two steps:
221 (i) images with any pixel with solar elevation lower than 10° are removed, leaving only images with all their
222 pixels in the daytime, and (ii) manual inspection over consecutive images with large mean differences, as
223 a way to detect images with corrupted pixels (bad acquisition or missing sectors). The first filter removes
224 images with any pixel at night but also ensures that the remaining images do not have pixels at sunrise and
225 sunset, which are moments of very low solar elevation over the horizon that appear distorted in the images
226 after the geometrical normalization with the cosine of the solar zenith angle (θ_z). The second filter separates
227 images for possible reconstruction. The separated images with less than 5% corrupted pixels were corrected
228 by inpainting. Specifically, the Navier-Stokes method of Bertalmio et al. (2001) was run. This method was
229 found to be the fastest of those in the OpenCV library. The percentage of images affected by the inpainting
230 was 1.2%, but only 0.12% had more than 1% of the pixels modified. Since less than 0.02% of the pixels
231 were inpainted a negligible impact is expected from the choice of the inpainting algorithm. After filtering
232 and reconstruction, the final data set is composed of 20842 valid images.

233 The images require a last preprocessing before their utilization. The original images contain reflectance
234 factor information (instead of Earth albedo), a dimensionless quantity that scales the measured radiance
235 recorder by the satellite’s radiometer by the maximum value that the sensor can detect. This includes the
236 correction to account for the variable Sun-Earth distance across the Earth’s orbit. This magnitude has a
237 dependence on the incidence angle of the sunlight to Earth, thus having geometrical spatial information.
238 Normalization is done by dividing each pixel by its corresponding $\cos(\theta_z)$ at each moment. The normalization
239 removes the deterministic geometrical variability coming from the apparent movement of the Sun relative to
240 the Earth and allows the comparison of images taken at different moments. All pixels exceeding the initial
241 range are clamped and the range of values is linearly mapped to $[0, 1]$. By doing this normalization, albedo
242 images are calculated from the original images and the data set is ready for utilization.

243 2.3. Data set split

244 The data set was divided into two subsets following (Nielsen et al., 2021b; Sønderby et al., 2020; Su
245 et al., 2020), one for training (75%) and the other for testing (25%). The random split was done day-wise
246 and distributed across the whole year. This resulted in 274 randomly selected days going into the training
247 split. The testing data set is then composed of 92 days that are unseen during the architecture definition
248 and parameters’ training. The validation data set, used for architecture decisions, included 40 random days
249 (15%) from the training set. The K-fold cross-validation methodology was not implemented due to its high
250 computational cost, being unfeasible for this type of satellite data set given the training times.

251 3. Metrics

252 The metrics used in this article are the usual ones in solar forecasting (Yang et al., 2020) and machine
 253 learning literature. These metrics can be used either for performance evaluation or as optimization targets.
 254 A short discussion of this topic will be presented in Subsection 5.1.3 for the U-Net architecture. For
 255 performance evaluation, the usual metrics in the solar forecasting field are favored, as they are directly
 256 related to the satellite albedo forecasting problem and were used in Aicardi et al. to bridge the gap between
 257 the forecast evaluation of both quantities. The basic common metrics are:

$$\begin{aligned}
 \text{Mean Bias Error} \quad \text{MBE} &= \frac{1}{N} \sum_{i=1}^H \sum_{j=1}^W (\hat{y}_{ij} - y_{ij}), \\
 \text{Mean Absolute Error} \quad \text{MAE} &= \frac{1}{N} \sum_{i=1}^H \sum_{j=1}^W |\hat{y}_{ij} - y_{ij}|, \\
 \text{Mean Squared Error} \quad \text{MSE} &= \frac{1}{N} \sum_{i=1}^H \sum_{j=1}^W (\hat{y}_{ij} - y_{ij})^2, \\
 \text{Root Mean Square Error} \quad \text{RMSE} &= \sqrt{\frac{1}{N} \sum_{i=1}^H \sum_{j=1}^W (\hat{y}_{ij} - y_{ij})^2},
 \end{aligned}$$

258 where \hat{y} and y are the predicted and ground truth images, respectively. Each image is of dimensions (H, W)
 259 and N denotes the number of valid pixels in the summation (which may be lower than $H \times W$). These
 260 metrics are calculated image-wise for each forecast horizon and can be expressed as a percentage of the
 261 image’s average value. These normalized metrics will be denoted with a % symbol. The performance
 262 evaluation is done by computing the average for each one of these metrics across all images in the test set,
 263 both for the normalized and not normalized cases. Note that RMSE is more sensitive to outliers than MAE
 264 (as it weighs more large variances) and that minimizing MAE or RMSE does not necessarily imply taking
 265 the MBE to zero.

266 An extended performance metric for solar irradiance (and meteorological) deterministic forecast is the
 267 forecasting skill, defined as:

$$\text{Forecasting Skill (\%)} \quad \text{FS} = 100 \times \left(1 - \frac{\text{RMSE}_{fcs}}{\text{RMSE}_{per}} \right),$$

268 where fcs refers to the forecast being evaluated and per refers to the persistence. This metric measures the
 269 gain, in terms of RMSE, of the forecast being evaluated relative to a baseline performance reference given
 270 by the persistence naive procedure. A positive (negative) metric means that the forecast is better (worse)
 271 than the baseline. The persistence procedure used in this work for the albedo satellite images is described

272 in [Subsection 4.1](#). All of the metrics are evaluated for each forecast horizon independently.

273 Finally, another metric that is used in this work is the Structural Similarity Index Measure (SSIM),
274 introduced by [Wang et al. \(2004\)](#). It measures the structural degradation between a distorted image and a
275 reference, thus it can quantify textures and perceptual similarity of an image relative to the reference. This is
276 a widely-adopted metric in the image processing community ([Fan et al., 2019](#)). However, it combines texture
277 and alignment in a single score, which was ultimately uninformative in most of the forecasting tests. Its fur-
278 ther utilization for this forecasting problem requires specific studies. This metric is introduced here because
279 it is used for an interesting study regarding the training metrics, which is presented in [Subsection 5.1.3](#).

280 4. Models

281 This section briefly describes the baseline models and the deep learning architectures considered in the
282 work. For more detailed information on the methods, the reader is referred to the original articles. The
283 experiments on the U-Net and IrradianceNet architectures are discussed in [Section 5](#). The CMV description,
284 implementation, and optimization are fully presented in this section.

285 4.1. Persistence

286 Regular persistence procedures use the last observation as the prediction. In this work, persistence is
287 implemented by simply maintaining the time t albedo image constant across all forecast horizons, providing
288 a simple baseline performance reference. Note that the aim of the work is not to assess to which extent
289 the models outperform a given reference, but to compare different models that are known to have a better
290 performance than naive forecasting procedures. In this sense, persistence provides a general reference and a
291 common ground to calculate the forecasting skill metric.

292 4.2. U-Net architecture

293 The U-Net ([Long et al., 2015](#); [Ronneberger et al., 2015](#)) is a well-known deep learning architecture for
294 semantic segmentation and pixel-wise prediction. It involves two stages, an encoder and a decoder in a
295 U-shaped scheme, both with several convolutional layers and skip connections between them. Through the
296 encoder, the network learns features and patterns in the image sequence. In the decoder, the prediction is
297 built by extracting and upsampling the features learned at different levels of the encoder’s convolutional lay-
298 ers. The skip connections help the decoder reconstruct the output by providing additional information, and
299 make the learning more stable by reducing the risk of vanishing gradients. The number of trainable param-
300 eters in this architecture is normally several million, whose adjustment can be done in modern computing
301 facilities equipped with Graphics Processing Units (GPUs). It has to be noted that the computationally
302 expensive part of the process is the training stage, both in terms of processing time and memory allocation,
303 but once the architecture is trained, its utilization can be performed in a few seconds without the need for

304 high computational resources. As this is an image-to-image deep learning technique, readers are referred to
305 the specific bibliography of the field for further details, which includes convolutional neural networks and
306 autoencoders. The architecture has been used on a plethora of problems (Falk et al., 2019; Wei et al., 2019;
307 Du et al., 2020; Smith et al., 2020; Kang et al., 2022) and has yielded a solid performance across a wide
308 variety of domains, from medical imaging to remote sensing.

309 The basic architecture has led to different variants of the U-Net. Three of them are here revisited,
310 namely, the Attention U-Net (Oktay et al., 2018), the Nested U-Net (Zhou et al., 2020), and the Recurrent
311 Residual U-Net (R2U-Net) (Alom et al., 2018). In the Attention U-Net, the main modification is to add the
312 so-called soft-attention layers, which allow the detection of the most relevant regions in the input images
313 to assign more importance to those while processing. The Nested U-Net modifies the skip connections by
314 adding more convolutional layers between the encoder and decoder, referred to as dense skip connections,
315 which are expected to reduce the semantic gap between the feature maps and the predictions. Finally,
316 the R2U-Net uses recurrent convolutional layers to feedback residuals in the training process instead of the
317 original convolutional layers. Further details on these variants can be found in their corresponding articles.
318 The use of these architectures is tested in [Subsection 5.1.1](#).

319 4.3. IrradianceNet

320 IrradianceNet is a custom convolutional long short-term memory (ConvLSTM) neural network-based
321 prediction model with a two-stage separated encoder-decoder scheme designed by Nielsen et al. (2021b) for
322 solar irradiance forecasting using geostationary satellite information. This method is reportedly the best-
323 performing of this family of forecasting techniques. It uses three ConvLSTM layers in both the encoder
324 and the decoder networks and employs a patch-based approach similar to Sønderby et al. (2020) due to
325 computational restrictions. The proposal uses four previous images as input to generate a prediction. Apart
326 from the satellite images, the authors introduced other sources of information as input: temporal information
327 as the hour, day, and month, and spatial information as the longitude, latitude, and elevation. A different
328 model is trained for each forecast horizon from 1 to 4 hours ahead. For the implementation of IrradianceNet
329 in this work, two versions of the model are considered, one with only satellite images and the other with
330 satellite images and geographic information, i.e. the longitude, latitude, and elevation. Full details on this
331 architecture are given in the article by Nielsen et al..

332 4.4. CMV

333 Cloud Motion Vectors (CMV) methods estimate the cloudiness velocity field from the last two consecutive
334 satellite images and then use it to generate future images, i.e. pixels are projected to their future position
335 by using the velocity field and the time t image. Several techniques have been applied to estimate the cloud
336 motion field, being the optical flow methods the most recent and best performing, as discussed in [Section 1](#).

337 In this work, due to ease of implementation and performance, the [Farneback](#) optical flow method is used
 338 with its OpenCV 3.x implementation (`calcOpticalFlowFarneback` function). This method requires some input
 339 parameters that were locally optimized over the training set. Some parameters refer to the mathematical
 340 formulation of the method, like the window size in which the polynomial expansion of this method is done.
 341 Other parameters refer to its computational implementation, like the down-scaling levels that are used to
 342 obtain the dense motion estimation from lower to higher resolution images, with a multi-level pyramid
 343 strategy. This method’s parameters and their optimized values are presented in [Table 1](#) with their library
 344 names. More information about this method can be found in [Farneback \(2003\)](#); [Aicardi et al. \(2022\)](#) and
 345 in its OpenCV documentation. It shall be noticed that the values of the parameters `winsize` and `levels` are
 346 similar to those found in [Aicardi et al.](#) for the same region, but with 2016-2017 GOES-13 albedo images, that
 347 have a different spatial resolution and time rate. This previous work only optimizes these two parameters,
 348 leaving the others as default.

Table 1: Optimized parameters for the [Farneback](#) optical flow method.

| parameter | <code>pyr_scale</code> | <code>levels</code> | <code>winsize</code> | <code>iterations</code> | <code>poly_n</code> | <code>poly_sigma</code> |
|------------------|------------------------|---------------------|----------------------|-------------------------|---------------------|-------------------------|
| value | 0.3987 | 4 | 22 | 3 | 5 | 0.8480 |

349 The projection algorithm being used is the common backward search, in which the predicted image is
 350 constructed by using the opposite vector flow at each given pixel and scaled by the time interval. Note
 351 that the scaling is required as the CMV is estimated with a 10-minutes difference between images and the
 352 forecast horizons are hourly. The scaled and inverted CMV is used to obtain the value to assign to each
 353 pixel in the predicted image from the previous image via a bi-linear interpolation. This procedure is iterated
 354 for all forecast horizons by taking the basis in the previous predicted image at each stage, starting with the
 355 time t real image. In [Aicardi et al.](#) it is shown that this iterative procedure is the best option to obtain the
 356 predicted images for these kinds of CMV algorithms.

357 [Aicardi et al.](#) also showed that running a spatial blurring on the predicted images improves the forecasting
 358 performance. Further, it showed that the blur window size should be horizon-dependent, as this provides
 359 better performance than a fixed spatial blur across all forecast horizons. The blurring implemented here
 360 is based on an isotropic Gaussian kernel. The size of this kernel was optimized over the training data to
 361 minimize the RMSE% between the predicted images and the corresponding ground truth for each forecast
 362 horizon. The analysis is shown in [Figure 2](#). As can be seen, there is a flat optimum value of kernel size for
 363 each forecast horizon that increases and flattens with increasing lead time. The behavior of this plot is the
 364 same as that of [Aicardi et al.](#), which uses a simple average value in a square spatial window of variable side
 365 length. This blurred model is the best-performing CMV strategy and is called here Blurred CMV.

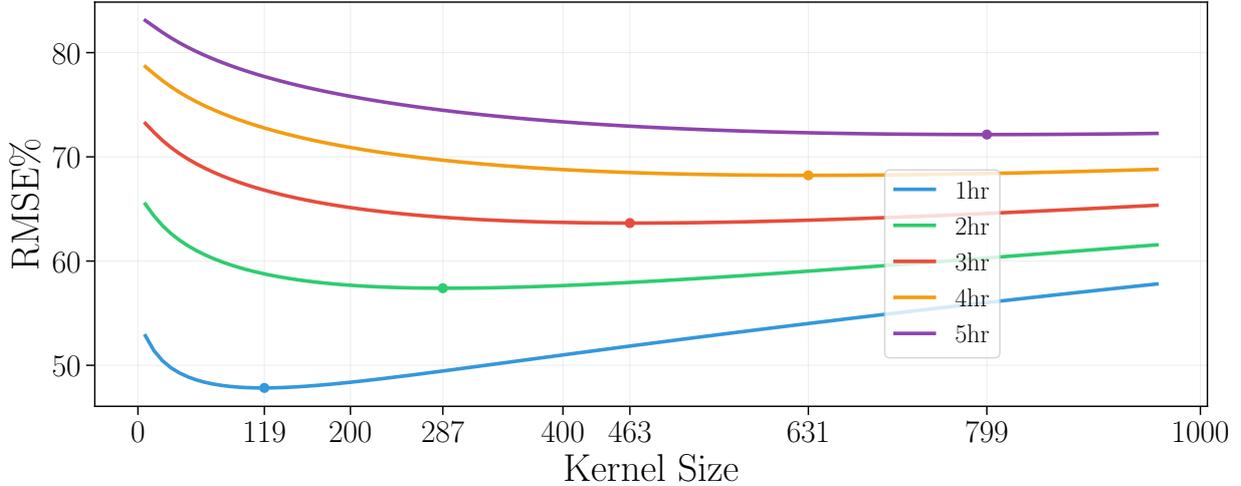


Figure 2: Kernel size optimization for the Blurred CMV strategy.

366 5. Ablation

367 This section describes several implementations of DL models, some of which yield promising results over
 368 the validation set. These preliminary best-performing architectures were selected for the final assessment
 369 with the test data set. The experiments presented in this section were obtained with the validation data
 370 set and a forecast horizon of 1 hour, favoring the metrics that are closer to the machine learning field. The
 371 red region of Figure 1a was used for most experiments, with the only exception of the variations of the
 372 U-Net architecture experiment (Subsection 5.1.1) that was performed on the smaller white bounding box
 373 of Figure 1a due to computational restrictions. All the experiments and runs required for this work were
 374 done in the ClusterUY center (Nesmachnow and Iturriaga, 2019), a national supercomputing infrastructure
 375 in Uruguay.

376 5.1. U-Net

377 Our proposed U-Net architecture uses 16 filters in the first layer (that will be simply referred to in
 378 the following as “filters”) and MAE as the training loss. Several variations were tested before arriving at
 379 this configuration. Experiments were done regarding the network capacity, the training loss function, and
 380 the architecture itself. The batch size was set as the maximum possible in each case. These experiments
 381 are described in different subsections in the following. This section also presents results regarding data
 382 augmentation and the inclusion of extra input information. Although not all of the experiments yielded
 383 positive improvements, they are briefly described here as they may be of interest to other researchers in the
 384 field. The baseline input information for the networks are the last three available images, this is, the images
 385 at times $(t - k)$ with $k = \{0, 1, 2\}$.

386 *5.1.1. U-Net variations*

387 The original U-Net and its variants, Attention U-Net, Nested U-Net, and R2U-Net, were compared to
 388 each other. These architectures were tested in the smaller-crop data set with 64 filters, the larger amount
 389 that was computationally possible. The variants (other than the original U-Net) have around 36 million
 390 trainable parameters, approximately twice the number of the largest U-Net considered in this work. This
 391 limits the maximum image size, directly related to the training time and required memory allocation.

392 The results of these tests are presented in [Table 2](#). It shows the training and validation MAE and the
 393 validation MSE (a proxy for the RMSE), and it is sorted by validation performance. It is observed that the
 394 performance of the variants at their maximum capacity does not improve the performance of the original
 395 U-Net, neither in training nor validation. It also shows that the U-Net variants tend to overfit more, as they
 396 have a poorer generalization performance relative to their training performance. This may suggest that the
 397 data set information is small relative to the networks’ capacity. In light of the previous analysis and the
 398 added complexity of the U-Net variants, only the original U-Net architecture was considered in the following
 399 tests of this section.

Table 2: Validation metrics for the different architecture variations of the original U-Net model with 64 filters. The test was done for 1h ahead forecast over the smaller-crop data set and MAE as training loss.

| architecture | training MAE | validation MAE | validation MSE |
|---------------------|---------------------|-----------------------|-----------------------|
| original U-Net | 0.0555 | 0.0780 | 0.0182 |
| attention U-Net | 0.0597 | 0.0870 | 0.0198 |
| nested U-Net | 0.0746 | 0.0931 | 0.0221 |
| R2U-Net | 0.0614 | 0.1053 | 0.0283 |

400 *5.1.2. Number of parameters*

401 In the original U-Net architecture the number of parameters is directly proportional to the number of
 402 filters. Three networks with different numbers of filters (16, 32, and 64) were trained for a prediction horizon
 403 of one hour. The first two rows of [Table 3](#) show the configurations that were tested, in which the relation
 404 between the number of filters and parameters can be observed. From these tests, it can be concluded that
 405 the U-Net with 16 filters achieves better training and validation performance than the U-Nets with 32 and
 406 64 filters. Note that the metrics of [Table 2](#) and [Table 3](#) do not match exactly due to different images’ sizes
 407 and coverage, although their behavior and order of magnitude are the same. The U-Net configuration with
 408 16 filters is the baseline U-Net used in the next experiments.

Table 3: Validation metrics for the U-Net model with different number of filters and 1h ahead forecast.

| filters | parameters | training MAE | validation MAE | validation MSE |
|---------|------------|--------------|----------------|----------------|
| 16 | 1080929 | 0.0532 | 0.0631 | 0.0119 |
| 32 | 4318401 | 0.0591 | 0.0766 | 0.0171 |
| 64 | 17262977 | 0.0587 | 0.0920 | 0.0238 |

409 *5.1.3. Impact of the training metric*

410 When training a machine learning model, an optimization objective must be chosen, also known as loss
 411 function. Naturally, a model trained to minimize a specific metric can perform sub-optimally under another
 412 metric (Zhao et al., 2017). Which one of the metrics is the most adequate for the solar forecasting problem
 413 is still an open question, as the value of the forecast is directly related to its ability to influence the decision-
 414 making processes (Yang et al., 2022). However, there is agreement that a set of independent metrics can
 415 assess different aspects of the quality of a forecast (Yang et al., 2020) if chosen carefully. A related question is
 416 which should be the training metric for models aimed at cloudiness and solar deterministic forecast, with the
 417 added complexity of spatial representation, which is different from the single-location time-series analysis.
 418 An analysis in this sense is provided in this subsection, inspecting mainly the MAE and MSE as training
 419 metrics, to understand the effect of the loss function choice.

420 Three U-Net architectures with 16 filters were trained with the only difference of its training loss function,
 421 being respectively, the MAE, MSE, and SSIM. The test was done in the same conditions as the previous
 422 subsection (for one hour ahead forecast with the 1024×1024 px images). MAE and MSE were selected as
 423 they represent classic metrics in the solar forecast field that are typically used for the adjustment of machine
 424 learning methods. These metrics weigh the forecast errors differently and quantify different aspects of the
 425 forecast quality. SSIM metric was also included as an image quality metric, mainly as an exploratory option.
 426 This metric has a quite different conception and objective than the MAE and MSE. The performance of
 427 the three models was assessed over the validation set with the same three metrics, obtaining the double-
 428 entrance 3×3 matrices of Table 4. The left-hand side of this table presents the three validation metrics (in
 429 absolute terms) when each of them is used as a loss function for training. The right-hand side shows the
 430 same information but centered by subtracting the optimum value of each metric across the three tests (the
 431 minimum for MAE and MSE, and the maximum for SSIM) and expressed as a percentage of it. Please note
 432 that MAE and MSE are negative-oriented metrics while SSIM is positive-oriented. As expected, each model
 433 performs better when evaluated with the same metric used for training. It is observed that both MAE and
 434 MSE can be used as loss functions without much loss in the other metrics. MAE as the optimization target
 435 has slightly less impact on the SSIM than the MSE. On the other hand, using MAE as the loss function
 436 degrades the MSE more (8.2%) than the reverse situation, in which using MSE degrades MAE by 4.1%.

437 The use of the SSIM as the loss function increases importantly the MSE validation metric and has less
438 impact on the MAE validation metric. This is consistent with the sharper visual results of the MAE in
439 comparison to the MSE, as this latter tends to generate predicted images with higher blur. All in all, as the
440 choice of MAE or MSE as the loss function has a different relative impact (higher or lower) in the other two
441 metrics, the analysis is not conclusive. The MAE metric was then favored as it is common ground in the
442 machine learning field. If the RMSE is chosen as the evaluation metric, one may expect an improvement of
443 $\simeq 4\%$ ¹ from the results presented in this article (Section 6) by using instead the MSE as loss function for
444 the U-Nets’ training.

Table 4: Validation metrics for the U-Net model and 1h ahead forecast when trained using three different loss functions.

| validation metric | cost function | | | percentage difference | cost function | | |
|----------------------|---------------|---------------|---------------|--------------------------|---------------|-------------|-------------|
| | MAE | MSE | SSIM | | MAE | MSE | SSIM |
| MAE (abs.) | 0.0631 | 0.0657 | 0.0669 | MAE ($\Delta\%$) | 0.0% | 4.1% | 6.0% |
| MSE (abs.) | 0.0132 | 0.0122 | 0.0153 | MSE ($\Delta\%$) | 8.2% | 0.0% | 25.4% |
| SSIM (abs.) | 0.5905 | 0.5838 | 0.5979 | SSIM ($\Delta\%$) | -1.2% | -2.4% | 0.0% |

445 The previous analysis provides a first study on the relationship between the training loss function and
446 the target performance metrics, an issue that has not been extensively addressed in the field so far. Of
447 course, if one particular metric is of interest for whatever problem-specific reason, then it should be used as
448 the optimization target for the training stage. An interesting discussion of one part of this problem can be
449 found in Section 2.1 of Yang et al. (2020).

450 5.1.4. U-Net Diff

451 The U-Net Diff model consists of changing the target to be the difference between the last image (time
452 t) of the input sequence and the desired objective for the given time horizon. In this way, the U-Net Diff
453 is trained to predict the changes between the actual and the future image, and not the future image itself.
454 This can be seen as a naive way to remove the image’s background. To allow for negative values in the
455 output, the last activation function was changed to a hyperbolic tangent instead of a sigmoid. To test this
456 modification, the baseline U-Net and the U-Net Diff were trained for prediction horizons from 1 to 5 hours
457 ahead with the 1024×1024 px images using MAE as the loss function. The validation results show similar
458 performance for the two networks, with the U-Net Diff achieving marginally better results in four of the five
459 forecast horizons. As the results over the validation set are very similar to the results over the final test set,
460 only the latter are presented in Section 6.

¹This was assessed from Table 4 by calculating $RMSE = \sqrt{MSE}$.

461 5.1.5. Extra ablations

462 This subsection summarizes some other tests that caused no improvement but might be of interest. The
463 preliminary studies about the optimal input data pipeline showed that no performance improvement came
464 from (i) using rotations as a data augmentation strategy (a strategy that is used in some contexts), (ii)
465 adding the date and time as input information for the network, or (iii) using the CMV method as extra
466 input information. The CMV information was included in the form of forecasts or vector fields, either
467 as additional channels or at a later network stage, and none of them produced improvements. Another
468 experiment regarding implementation showed that recursively using a 10-min single-horizon U-Net resulted
469 in worse performance than using the horizon-specific models. This was the expected behavior, as the U-Net
470 with a 10-min horizon was not trained to deal recursively with its blurry outputs as inputs or to minimize
471 the recursive error on larger horizons, but was tested for completeness.

472 5.2. IrradianceNet

473 Two types of IrradianceNet models were trained. One only uses images as input and the other the images
474 with added spatial (latitude and longitude) and topographic data (elevation map). The elevation map was
475 normalized by its absolute maximum inside the corresponding crop. The other spatial inputs (coordinates)
476 were mapped to the $[0, 1]$ interval. The network using these extra inputs is called IrradianceNet GEO and was
477 recommended by [Nielsen et al. \(2021b\)](#). As it is trained over patches of the images due to its computational
478 cost, the spatial information may provide the network with knowledge from where the patch is taken. A
479 10-minute prediction step was too computationally expensive, so a 30-minute time step was used instead,
480 following the original article by [Nielsen et al.](#) Also, to be consistent with the original article, both models
481 use four previous images as input to generate the prediction (times $(t - k)$ with $k = \{0, 1, 2, 3\}$), although
482 in this case, they are separated by 10-minute steps. Note that, in this way, this algorithm uses one more
483 previous image than the U-Nets. A related issue is that, operationally, the algorithm needs to wait for one
484 daylight image more than the versions with three inputs, delaying 10 minutes (in this case) its first forecast
485 of the day. This is not an important cost when using 10-minute images, but it certainly would be an issue
486 with 30-minute images.

487 The training configuration for IrradianceNet and IrradianceNet GEO is similar to the ones used for the
488 U-Net: 100 epochs, Xavier initialization ([Glorot and Bengio, 2010](#)) for the weights, and Adam optimizer
489 ([Kingma and Ba, 2015](#)) with a variable learning rate scheduled to be reduced in half if validation perfor-
490 mance does not improve for 15 epochs. However, there are two changes to be consistent with the training
491 configuration used in the original article. These are the initial learning rate equal to 2×10^{-3} and the
492 selection of MSE as the training loss. The location of the patches taken during training is random for each
493 batch. During validation, the patches are fixed to cover the whole image without overlapping.

494 The results from running the trained models on the validation data set showed that the IrradianceNet

495 GEO was not able to exploit the additional input data, as it performed very similarly to the basic Irradi-
496 anceNet. These results could be explained by the low value that the geographic information adds for the
497 South American Pampa Húmeda (which is mostly flat grassland) in comparison to the region of the original
498 article (Europe). IrradianceNet GEO is included in the final results for completeness.

499 6. Results

500 This section contains the final results of the performance evaluation and the method’s behavior over
501 the test data set, considering the proposed and selected models. This test data set has been unseen in all
502 previous analyses and optimization. [Subsection 6.1](#) presents the final performance assessment and metrics,
503 while [Subsection 6.2](#) shows some selected predicted sequences with interesting insights. For clarity, some of
504 the quantitative information is presented in [Appendix B](#).

505 6.1. Quantitative results

506 [Figure 3](#) shows the MAE, RMSE, and FS metrics across the hourly forecast horizons up to 5h ahead. The
507 evaluation includes three reference algorithms: the Persistence and the CMV in its two versions, pixel-wise
508 and with optimized spatial blurring. The two pairs of DL models being evaluated are included, the U-Net
509 and U-Net Diff, and the IrradianceNet with and without spatial information. As explained in [Section 1](#), the
510 CMV algorithm is evaluated only over its valid pixels. For the rest of the models, the evaluation is conducted
511 over the entire 1024×1024 px images. Please note that the metrics’ values do not need to match those
512 of [Section 5](#) as the data sets are not the same (validation vs test data set). However, it is a sanity check
513 to observe that the orders of magnitudes are the same. A direct qualitative comparison with the results of
514 [Nielsen et al. \(2021a\)](#) or [Nielsen et al. \(2021b\)](#) over Europe is not feasible due to different reasons. In [Nielsen](#)
515 [et al. \(2021a\)](#) the algorithm is run with a satellite cloud classification product, so the reported metrics at the
516 image level are in accordance with the objectives of the study, being different from the ones here. In [Nielsen](#)
517 [et al. \(2021b\)](#) the evaluation is performed for the hourly predicted irradiation, not at image level (which
518 would be in that case for the effective cloud albedo, a satellite cloud index used for satellite-to-irradiance
519 conversion), and, apart from MAE, the metrics are different. In any case, IrradianceNet is included in this
520 article and evaluated with the same data set along with the other methods, thus providing a fair comparison
521 in the target region. The performance differences that may arise due to different regions and climates should
522 be addressed in future benchmark studies.

523 The largest difference between performances in [Figure 3](#) is seen when evaluating with MAE, where the
524 U-Net and U-Net Diff significantly outperform all the other models. This is expected as these are the only
525 two models trained to optimize this metric. It can be seen that the two IrradianceNet versions have almost
526 the same performance in this metric and outperform the persistence and pixel-wise CMV. In comparison

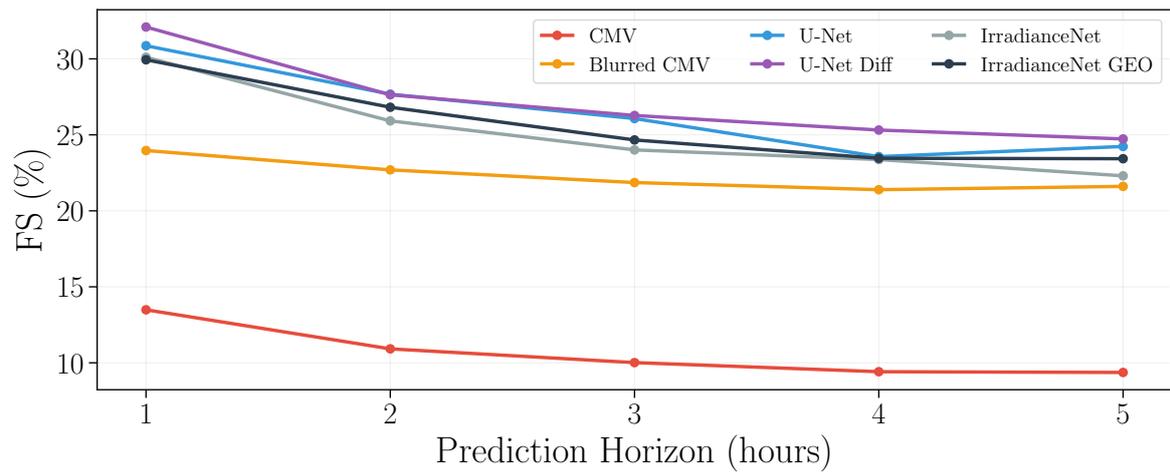
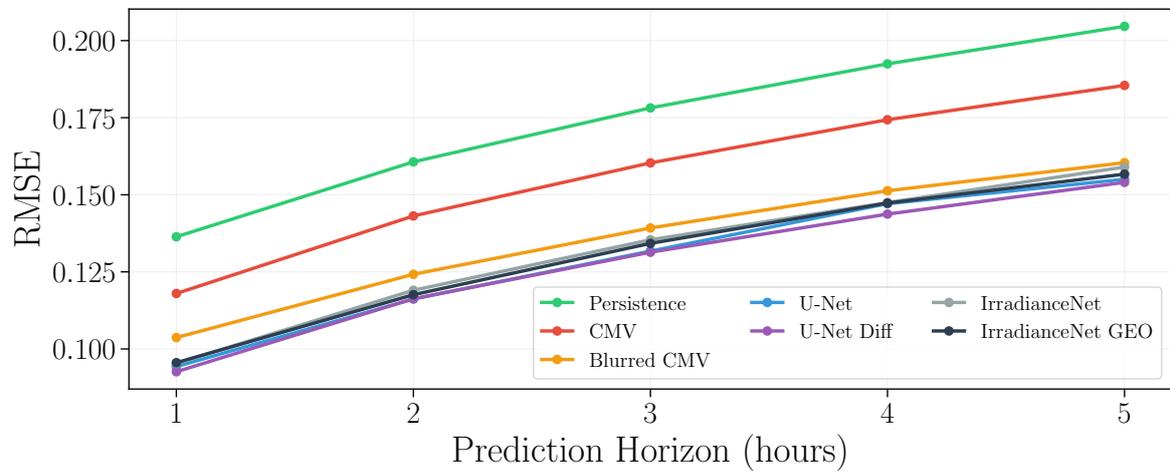
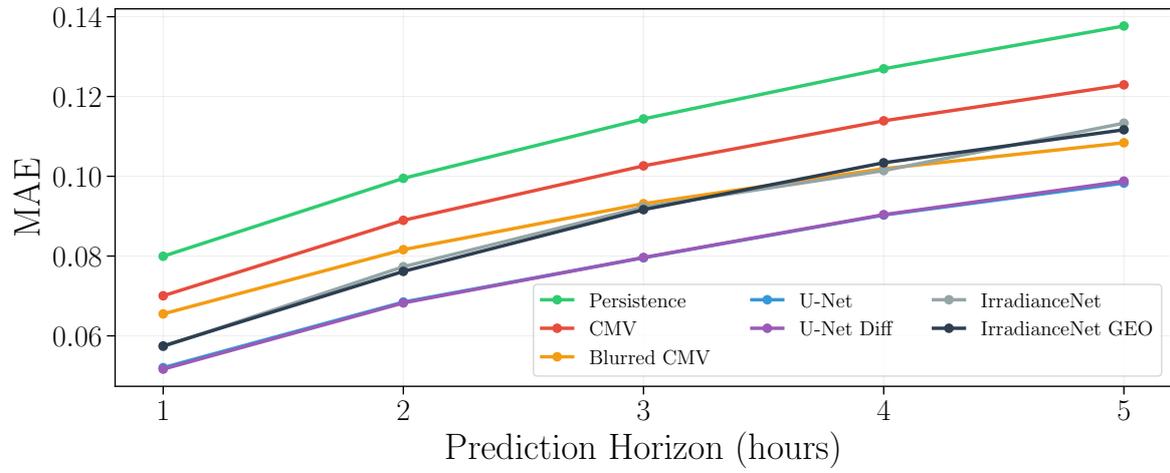


Figure 3: Main performance metrics vs forecast horizon for the reference methods and DL models.

527 with the blurred CMV, the IrradianceNet performance is similar for 3 and 4 hours ahead, being better in
528 the 1h and 2h forecast horizons and worse in the last one. It shall be recalled at this point that both the
529 Blurred CMV and IrradianceNet networks were optimized to minimize the training MSE.

530 Observing the RMSE performance plot of [Figure 3](#) the differences are smaller than in the MAE analysis,
531 and they are better observed in the Forecasting Skill plot. The IrradianceNet variations perform better
532 than the blurred CMV for all forecast horizons under this metric. The best-performing architectures are the
533 U-Nets, being the U-Net Diff the one with the highest FS% for all forecast horizons. The fact that the U-Net
534 is still superior in RMSE when using a different training loss is remarkable, as per [Table 4](#) this performance
535 could be further improved in about 4% if MSE were to be used as a loss function. The Forecasting Skill
536 plot also allows seeing the important improvement that is gained with a simple spatial blurring in the CMV
537 output, coinciding with the results of [Aicardi et al. \(2022\)](#) with the former GOES-East satellite images. The
538 work of [Aicardi et al.](#) also obtained a decreasing trend in the forecasting skill with the forecast horizon with
539 a higher drop in the transition between the first and the second hour, and with similar values. The MBE%
540 and RMSE% plots are shown and discussed in [Appendix B](#), and also have similar values to that of [Aicardi](#)
541 [et al.](#).

542 6.2. Qualitative results

543 It is interesting to explore some differences between the U-Nets, IrradianceNet, and CMV predictions.
544 Two comparisons are made: (i) the U-Net vs CMV, showing some selected cases in which the U-Net was
545 able to predict cloud extinction, a feature that current baseline CMV strategies are unable to perform, and
546 (ii) the spatial distribution of the errors of the U-Net and IrradianceNet networks, specifically addressing
547 the artifacts observed by using the patch processing strategy. Point (i) also includes a visualization of the
548 border's effect of the CMV methods. The U-Net Diff is selected for the following discussion.

549 6.2.1. Cloud extinction case studies

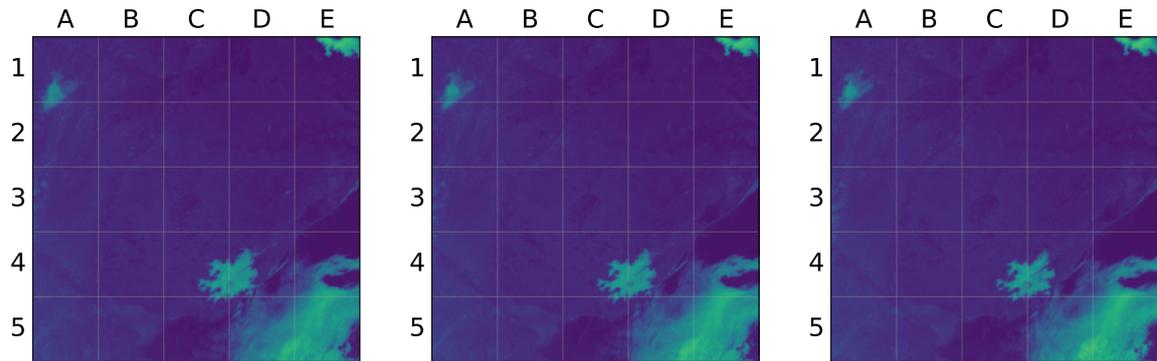
550 [Figures 4](#) and [5](#) show two examples in which the U-Net was able to correctly forecast cloud extinction,
551 while the CMV, as expected, was not. The first row shows the three lagged images used for the U-Net
552 forecast, and the one on the right is the time t image. Note that the CMV only uses the last two of these
553 three images. The second and third rows show the predictions of the U-Net and the CMV along with the
554 corresponding ground truth image on the left side. In the first sequence of [Figure 4](#) (August 4th, 2020,
555 time t being 14:40 UTC-0) the clouds in 4C-4D, 1A-2A, and 1E cells disappear after 2 hours, as can be
556 seen in the ground truth. The U-Net forecasts these extinctions accurately in the three cases, including the
557 intermediate stage at one hour ahead, being a remarkable feature of this method. The CMV, in change,
558 maintains these clouds in its prediction. The second sequence of [Figure 5](#) (October 14th, 2020, time t being
559 12:40 UTC-0) shows a similar behavior but in a more complex situation. In this case, the extinction of the

560 clouds located in cells 2D-2E and 4D occurs after three hours. The U-Net, again, manages to predict its
561 reduction and later disappearance, but in this case in a moving context. In particular, the U-Net detects the
562 drift of the 2D-2E clouds and it is capable of predicting that these clouds will detach from the main system
563 in the top-left corner, and gradually vanish for larger horizons. Similar behavior and accurate prediction
564 by the U-Net are observed for the cloud in the 4D cell. Of course, the CMV is unable to predict any of
565 these phenomena, and these identified clouds wrongly remain in all its predictions. This second sequence
566 also allows seeing the border's effect of the CMV prediction, as this sequence includes important movement
567 in the scene. This effect increases with the time horizon, being a drawback of the CMV strategy that can
568 be mitigated by using larger images, with the corresponding computational cost.

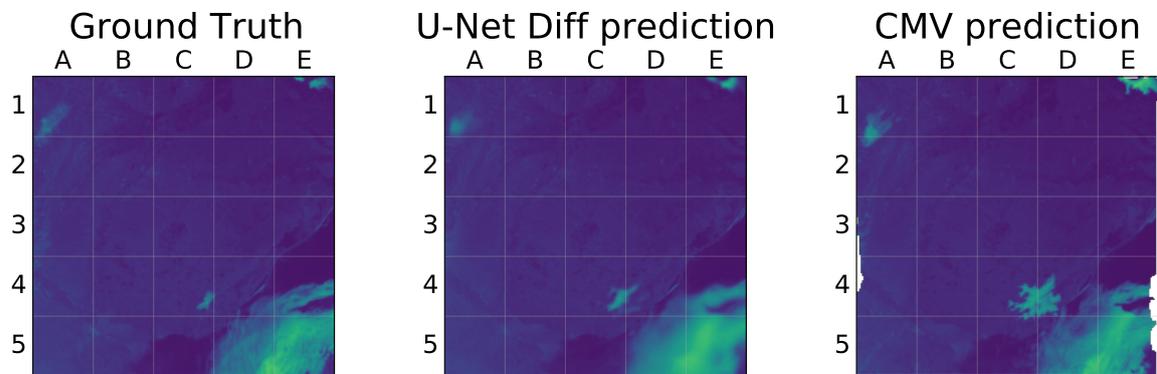
569 The previous features of the U-Net do not come along without any cost. As can be seen in both
570 sequences the U-Net predictions tend to be blurred, a trick that is learned by the network to reduce the cost
571 of high errors. As in any deterministic forecasting problem, there is a trade-off between having an overall
572 good performance and taking risks in the prediction. There is then important room for further studies
573 to understand the separation between image blurring and predicting clouds' movement, deformation, and
574 formation/extinction, some of which are part of our current work.

575 6.2.2. *Spatial dependence of the error*

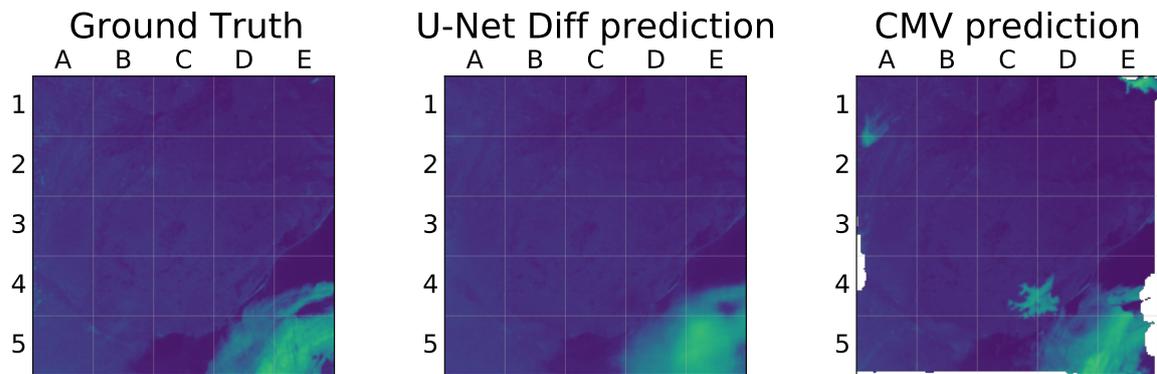
576 It is expected that the models perform better close to the center of the image, as there is less uncertainty
577 about the clouds that could move into or out of that area. In change, the prediction of incoming clouds
578 through the image borders is a much harder task. However, this may not be true, if clouds are more difficult
579 to predict in a given sub-region. Also, predicting the borders of the image is not a symmetric problem in
580 average terms, as the atmosphere circulation makes clouds' phenomena have preferred directions in different
581 parts of a given territory. This analysis was done with three models: the Persistence, the U-Net Diff, and
582 IrradianceNet methods. [Figure 6](#) presents the spatial per-pixel distribution of the RMSE over the test data
583 set for each of the 5 hourly time horizons. These error maps show that all models make, on average, the
584 largest errors in the top right corner (northeast) of the region under study, being the Persistence procedure
585 the most disadvantaged. This analysis reveals that the climate and geographical characteristics of a region
586 can be an important driver of errors' spatial distribution. For this region, in particular, the clouds in
587 the northeast region are harder to predict, being associated with the typical circulation and behavior of
588 cloudiness in that area. These areas with the highest errors are consistent through the 5-time horizons, with
589 the difference of a natural error increase in the whole map as the prediction horizon grows. When visualizing
590 the error maps of IrradianceNet, the patch-based prediction is observed as a grid-like layout throughout the
591 region. Apart from the grid-like effect, the underlying error maps of the IrradianceNet predictions are similar
592 to the ones of the U-Net Diff.



(a) Last three images (times $(t - k)$ with $k = \{0, 1, 2\}$), from left (14:20 UTC0) to right (14:40 UTC0).

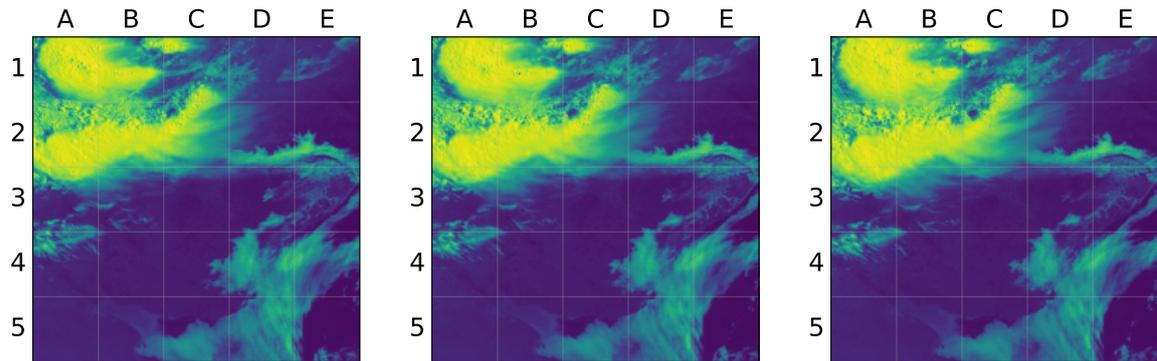


(b) 1 hour ahead ground truth and predictions.

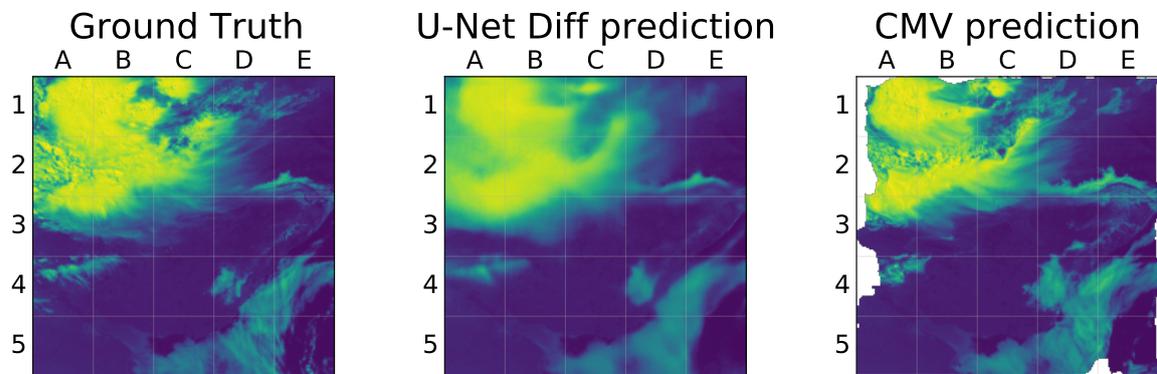


(c) 2 hours ahead ground truth and predictions.

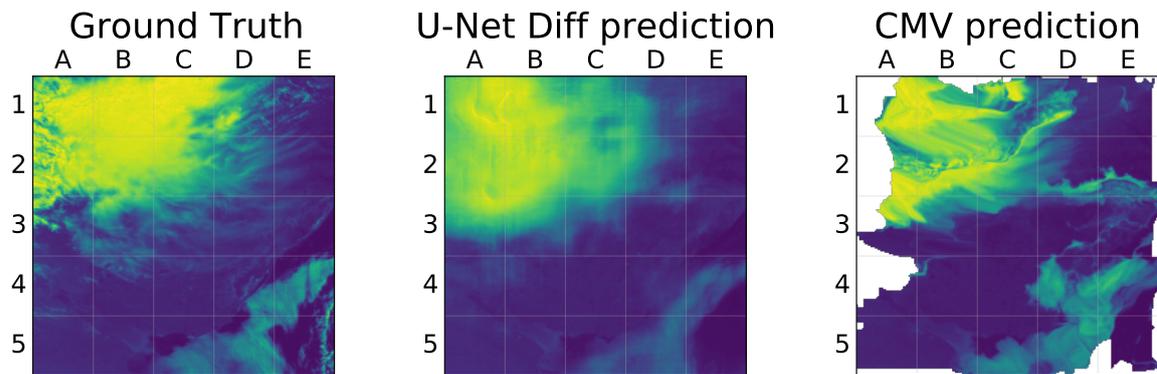
Figure 4: Image sequence captured on August 4th, 2020. Current time: 14:40 UTC0 (top-right image).



(a) Last three images (times $(t - k)$ with $k = \{0, 1, 2\}$), from left (12:20 UTC0) to right (12:40 UTC0).



(b) 1 hour ahead ground truth and predictions.



(c) 3 hours ahead ground truth and predictions.

Figure 5: Image sequence captured on October 14th, 2020. Current time: 12:40 UTC0 (top-right image).

593 7. Conclusions

594 This article analyzed Deep Learning (DL) techniques applied to satellite-based cloudiness prediction
 595 (Earth's albedo) up to 5 hours ahead. This is the first stage of satellite intra-day solar irradiance forecasting.

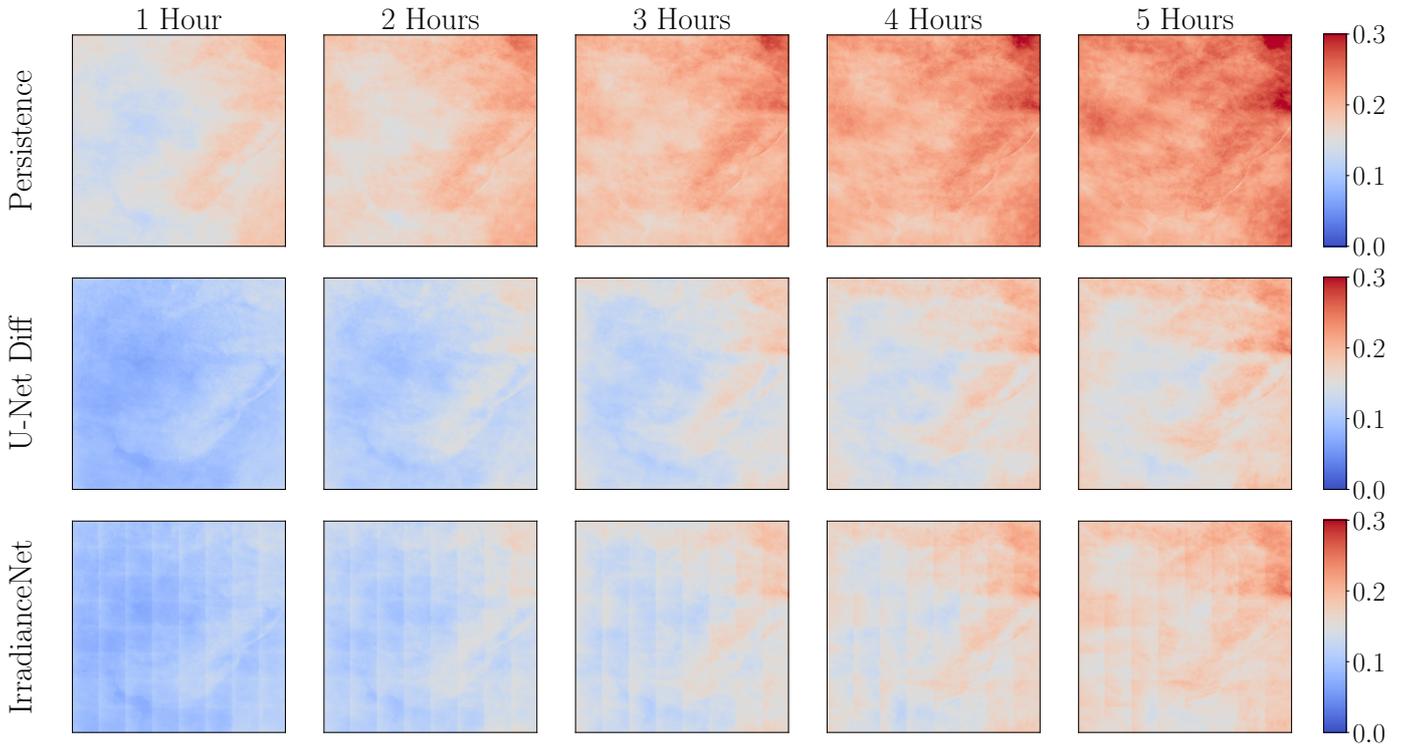


Figure 6: Root mean squared error per pixel for all five-time horizons.

596 The work used the 10-minute GOES-16 visible channel satellite images for the southeast part of South
597 America, an area known as Pampa Húmeda. This is a region in which the convective systems' evolution
598 is challenging and the solar irradiance variability is at an intermediate level, as clear sky, overcast, and
599 partly cloudy conditions alternate. The utilization of the original U-Net DL network was proposed and
600 tested for this purpose along with the up-to-date IrradianceNet DL algorithm. Two baseline methods were
601 also included, namely, the satellite cloudiness persistence and an advanced Cloud Motion Vectors (CMV)
602 strategy with optimized spatial blurring. The U-Net optimization was analyzed in detail, providing for it
603 different ablation studies. Two U-Nets with 16 filters were implemented; the regular U-Net that aims to
604 predict the next image and the U-Net Diff that aims to predict the difference with the last available image
605 (time t). A different network was trained for each time horizon. This is a common practice in the field,
606 however, it was confirmed in this work as best practice, in opposition to the recurrent utilization of a single
607 one-lead-time DL network. None of the U-Net variants (Attention U-Net, Nested U-Net, and R2U-Net)
608 were found to outperform the original U-Net. Both final U-Nets showed the remarkable feature of predicting
609 cloud extinction, which is one of the harder issues in satellite-based solar forecasting.

610 The DL methodologies presented better performance than the baseline methods, including the blurred
611 CMV, which sets a very exigent performance bound. The preexisting architecture, IrradianceNet, as pro-
612 posed by [Nielsen et al.](#), was adapted successfully when retrained and evaluated on a different geographical

613 region and larger images. In this study, at least for this region, both U-Net architectures outperformed Ir-
614 radianceNet, and the U-Net Diff was the best performing. One of the strengths of the U-Net architecture is
615 its wide utilization and ease of implementation, as it is considered a light, robust, and extensively tested DL
616 method. This work illustrated the U-Net utilization over images with a larger size and a higher resolution
617 than in previous works dealing with DL methods for this purpose. Regarding training time, the Irradi-
618 anceNet had higher requirements than the U-Nets. However, once the DL networks are trained, prediction
619 times are low, making any of them suitable for real-time operation. In particular, none of the final U-Net
620 models are computationally expensive to use, being possible to generate a prediction in less than a second
621 using a single GPU.

622 A. Limitations

623 There is plenty of room for improvement and this work is far from comprehensive. For instance, data
624 augmentation via random cropping was not considered for the U-Net. This method was used to train
625 IrradianceNet and can help escalate the models to larger regions (Espeholt et al., 2021). In addition, the
626 optimization objective remains arbitrarily defined, for there is not a single metric of interest. This introduces
627 optimization compromises, e.g. when the optimization is done with the MSE loss function, the predictions
628 appear with higher blur than when using MAE. There is still an unsolved (and not yet fully understood)
629 trade-off between overall accuracy and risky variability prediction, that the networks learn to mitigate by
630 blurring. Moreover, the preprocessing could also be enhanced by including image background removal in
631 sophisticated ways. This could be especially impactful for salty/snowy regions of land, not present in our
632 images. The impact of such a transformation on performance is yet unknown and it was not specifically
633 addressed in this work, being part of our current work. Lastly, the results suggested no improvement when
634 adding the GEO information to the IrradianceNet. Although this can be explained by the low variability in
635 altitude in the studied area compared to Europe, the question of what information can be effectively fed to
636 the neural network is raised. Another way to include meaningful information would be to directly provide
637 the neural network with relevant regional data, such as typical cloudiness variability or mean wind direction.

638 B. Complementary assessment information

639 For completeness and clarity of presentation in the main text, the MBE% and RMSE% plots are provided
640 here. Figure B.7 shows these two metrics for the assessment of Section 6. The MBE% are between $\simeq \pm 10\%$,
641 similar to Aicardi et al. (2022), in which all tested CMV strategies and the persistence tend to decrease (with
642 its sign) with the forecast horizons. In particular, persistence shows very similar behavior. The differences
643 here are observed with the IrradianceNet variations, whose MBE% increases with the forecast horizons,
644 being positive for 3 to 5 hours ahead. The RMSE% plot contains the same information as the RMSE plot

645 of Figure 3 but is normalized by each image’s mean value before averaging the metric across all images in
 646 the test set. It is interesting to note that the order of magnitude in this plot is similar to that of [Aicardi](#)
 647 [et al.](#).

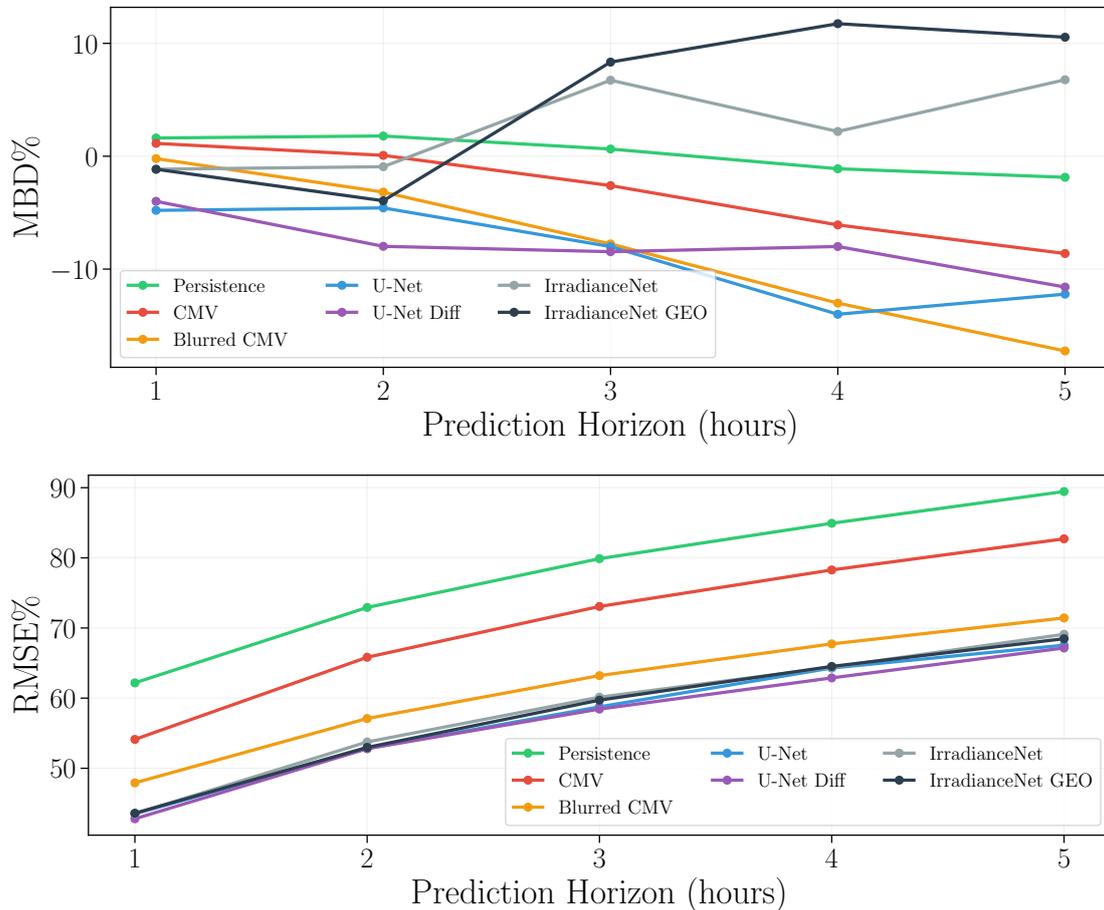


Figure B.7: Complementary performance metrics (%) vs forecast horizon for the reference methods and DL models.

648 Author contributions

649 **Franco Marchesoni-Acland** contributed with conceptualization, methodology development, writing
 650 - original draft, writing - review and editing, supervision, and project administration. **Andrés Herrera,**
 651 **Franco Mozo, and Ignacio Camiruaga** contributed with conceptualization, methodology development,
 652 software, validation, formal analysis, investigation, resources, data curation, writing - original draft, writing
 653 - review and editing, and visualization. **Alberto Castro** contributed with resources, writing - review and
 654 editing, supervision, and project administration. **Rodrigo Alonso-Suárez** contributed with conceptualiza-
 655 tion, methodology, resources, writing - original draft, writing - review and editing, supervision, and project

656 administration.

657 Acknowledgments

658 Data was made available by the Solar Energy Laboratory (LES) of the Universidad de la República. The
659 experiments presented in this article were performed using the ClusterUY infrastructure (<https://cluster.uy>).
660 R. Alonso-Suárez acknowledges partial financial support from the CSIC Group’s Program, Udelar, Uruguay.

661 References

- 662 Aicardi, D., Musé, P., and Alonso-Suárez, R. (2022). A comparison of satellite cloud motion vectors techniques to forecast
663 intra-day hourly solar global horizontal irradiation. *Solar Energy*, 233:46–60.
- 664 Alom, M. Z., Hasan, M., Yakopcic, C., Taha, T. M., and Asari, V. K. (2018). Recurrent residual convolutional neural network
665 based on U-Net (R2U-Net) for medical image segmentation. *CoRR*, abs/1802.06955.
- 666 Alonso-Suárez, R., David, M., Branco, V., and Lauret, P. (2020). Intra-day solar probabilistic forecasts including local short-
667 term variability and satellite information. *Renewable Energy*, 158:554–573.
- 668 Alonso-Suárez, R., Marchesoni, F., Dovat, L., and Laguarda, A. (2021). Satellite-based operational solar irradiance forecast
669 for Uruguay’s solar power plants. In *2021 IEEE URUCON*, pages 182–187.
- 670 Baranchuk, D., Rubachev, I., Voynov, A., Khrulkov, V., and Babenko, A. (2022). Label-efficient semantic segmentation with
671 diffusion models. Published as a conference paper at ICLR 2022, arXiv:2112.03126.
- 672 Bertalmio, M., Bertozzi, A., and Sapiro, G. (2001). Navier-Stokes, fluid dynamics, and image and video inpainting. In
673 *Proceedings of the 2001 IEEE Computer Society Conference on Computer Vision and Pattern Recognition. CVPR 2001*,
674 volume 1, pages I–I.
- 675 Berthomier, L., Pradel, B., and Perez, L. (2020). Cloud cover nowcasting with deep learning. In *Proceedings of the 2020 Tenth
676 International Conference on Image Processing Theory, Tools and Applications (IPTA)*, pages 1–6.
- 677 Brousseau, P., Seity, Y., Ricard, D., and Léger, J. (2016). Improvement of the forecast of convective activity from the AROME-
678 France system. *Quarterly Journal of the Royal Meteorological Society*, 142(699):2231–2243.
- 679 Cano, D., Monget, J., Albuissou, M., Guillard, H., Regas, N., and Wald, L. (1986). A method for the determination of the
680 global solar radiation from meteorological satellite data. *Solar Energy*, 37(1):31–39.
- 681 Croitoru, F.-A., Hondru, V., Ionescu, R. T., and Shah, M. (2023). Diffusion models in vision: A survey. *IEEE Transactions
682 on Pattern Analysis and Machine Intelligence*, pages 1–20. arXiv:2209.04747.
- 683 Cros, S., Sébastien, N., Liandrat, O., and Schmutz, N. (2014). Cloud pattern prediction from geostationary meteorological
684 satellite images for solar energy forecasting. In *Proceedings of SPIE - The International Society for Optical Engineering*,
685 volume 9242.
- 686 Du, G., Cao, X., Liang, J., Chen, X., and Zhan, Y. (2020). Medical image segmentation based on u-net: A review. *Journal of
687 Imaging Science and Technology*, 64(2):020508–1–020508–12.
- 688 Espenholt, L., Agrawal, S., Sønderby, C. K., Kumar, M., Heek, J., Bromberg, C., Gazen, C., Hickey, J., Bell, A., and Kalch-
689 brenner, N. (2021). Skillful twelve hour precipitation forecasts using large context neural networks. *CoRR*, abs/2111.07470.
- 690 Falk, T., Mai, D., Bensch, R., Çiçek, O., Abdulkadir, A., Marrakchi, Y., Böhm, A., Deubner, J., Jäckel, Z., Seiwald, K.,
691 Dovzhenko, A., Tietz, O., Bosco, C. D., Walsh, S., Saltukoglu, D., Tay, T. L., Prinz, M., Palme, K., Simons, M., Diester,
692 I., Brox, T., and Ronneberger, O. (2019). U-Net: deep learning for cell counting, detection, and morphometry. *Nature
693 Methods*, 16:67–70.

694 Fan, L., Zhang, F., Fan, H., and Zhang, C. (2019). Brief review of image denoising techniques. *Visual Computing for Industry,*
695 *Biomedicine, and Art*, 2(1):1–12.

696 Farneback, G. (2003). Two-frame motion estimation based on polynomial expansion. In Bigun, J. and Gustavsson, T., editors,
697 *Image Analysis*, pages 363–370, Berlin, Heidelberg. Springer Berlin Heidelberg.

698 Glorot, X. and Bengio, Y. (2010). Understanding the difficulty of training deep feedforward neural networks. *Journal of*
699 *Machine Learning Research - Proceedings Track*, 9:249–256.

700 Horn, B. K. P. and Schunck, B. G. (1981). Determining optical flow. *Artif. Intell.*, 17(1-3):185–203.

701 IRENA (2021). Renewable energy statistics.

702 Kang, Z., Huang, Z., and Lu, C. (2022). Speech enhancement using u-net with compressed sensing. *Applied Sciences*, 12(9).

703 Kingma, D. P. and Ba, J. (2015). Adam: A method for stochastic optimization. In *Proceedings of the 3rd International*
704 *Conference on Learning Representations (ICLR)*.

705 Kühnert, J., Lorenz, E., and Heinemann, D. (2013). Chapter 11 - Satellite-based irradiance and power forecasting for the
706 german energy market. In Kleissl, J., editor, *Solar Energy Forecasting and Resource Assessment*, pages 267–297. Academic
707 Press, Boston.

708 Laguarda, A., Giacosa, G., Alonso-Suárez, R., and Abal, G. (2020). Performance of the site-adapted cams database and
709 locally adjusted cloud index models for estimating global solar horizontal irradiation over the pampa húmeda. *Solar Energy*,
710 199:295–307.

711 Lauret, P., Alonso-Suárez, R., Le Gal La Salle, J., and David, M. (2022). Solar forecasts based on the clear sky index or the
712 clearness index: Which is better? *Solar*, 2(4):432–444.

713 Lefèvre, M., Oumbe, A., Blanc, P., Espinar, B., Qu, Z., Wald, L., Homscheidt, M. S., and Arola, A. (2013). McClear: a new
714 model estimating downwelling solar radiation at ground level in clear-sky conditions. *Atmospheric Measurement Techniques*,
715 *European Geosciences Union*, 6:2403–2418.

716 Long, J., Shelhamer, E., and Darrell, T. (2015). Fully convolutional networks for semantic segmentation. In *2015 IEEE Con-*
717 *ference on Computer Vision and Pattern Recognition (CVPR)*, pages 3431–3440, Los Alamitos, CA, USA. IEEE Computer
718 Society.

719 Lorenz, E., Hammer, A., and Heinemann, D. (2004). Short term forecasting of solar radiation based on satellite data. In
720 *EUROSUN2004 (ISES Europe Solar Congress)*, pages 841–848, Freiburg, Germany.

721 Lucas, B. D. and Kanade, T. (1981). An iterative image registration technique with an application to stereo vision. In
722 *Proceedings of the 7th International Joint Conference on Artificial Intelligence - Volume 2, IJCAI’81*, pages 674–679, San
723 Francisco, CA, USA. Morgan Kaufmann Publishers Inc.

724 Nesmachnow, S. and Iturriaga, S. (2019). Cluster-UY: Collaborative Scientific High Performance Computing in Uruguay. In
725 Torres, M. and Klapp, J., editors, *Supercomputing*, pages 188–202, Cham. Springer International Publishing.

726 Nielsen, A. H., Iosifidis, A., and Karstoft, H. (2021a). CloudCast: A Satellite-Based Dataset and Baseline for Forecasting
727 Clouds. *IEEE Journal of Selected Topics in Applied Earth Observations and Remote Sensing*, 14:3485–3494.

728 Nielsen, A. H., Iosifidis, A., and Karstoft, H. (2021b). IrradianceNet: Spatiotemporal deep learning model for satellite-derived
729 solar irradiance short-term forecasting. *Solar Energy*, 228:659–669.

730 Oktay, O., Schlemper, J., Folgoc, L. L., Lee, M., Heinrich, M., Misawa, K., Mori, K., McDonagh, S., Hammerla, N. Y., Kainz,
731 B., Glocker, B., and Rueckert, D. (2018). Attention U-Net: Learning where to look for the pancreas. In *Medical Imaging*
732 *with Deep Learning*.

733 Peel, M. C., Finlayson, B. L., and McMahon, T. A. (2007). Updated world map of the köppen-geiger climate classification.
734 *Hydrology and Earth System Sciences Discussions*, 11:1633–1644.

735 Peng, Z., Yoo, S., Yu, D., and Huang, D. (2013). Solar irradiance forecast system based on geostationary satellite. In *2013*
736 *IEEE International Conference on Smart Grid Communications (SmartGridComm)*, pages 708–713.

- 737 Perez, R., Ineichen, P., Moore, K., Kmiecik, M., Chain, C., George, R., and Vignola, F. (2002). A new operational model for
738 satellite-derived irradiances: description and validation. *Solar Energy*, 73(5):307–317.
- 739 Perez, R., Kivalov, S., Schlemmer, J., Hemker, K., Renné, D., and Hoff, T. E. (2010). Validation of short and medium term
740 operational solar radiation forecasts in the US. *Solar Energy*, 84(12):2161–2172.
- 741 Qu, Z., Oumbe, A., Blanc, P., Espinar, B., Gesell, G., Gschwind, B., Klüser, L., Lefèvre, M., Saboret, L., Schroedter-
742 Homscheidt, M., and Wald, L. (2017). Fast radiative transfer parameterisation for assessing the surface solar irradiance:
743 The Heliosat-4 method. *Meteorologische Zeitschrift*, 26(1):33–57.
- 744 Rasmussen, K. L., Zuluaga, M. D., and Houze Jr., R. A. (2014). Severe convection and lightning in subtropical South America.
745 *Geophysical Research Letters*, 41(20):7359–7366.
- 746 Ravuri, S., Lenc, K., Willson, M., Kangin, D., Lam, R., Mirowski, P., Fitzsimons, M., Athanassiadou, M., Kashem, S., Madge,
747 S., et al. (2021). Skilful precipitation nowcasting using deep generative models of radar. *Nature*, 597(7878):672–677.
- 748 Ren, X., Li, X., Ren, K., Song, J., Xu, Z., Deng, K., and Wang, X. (2021). Deep learning-based weather prediction: A survey.
749 *Big Data Research*, 23:100178.
- 750 REN21 (2021). Renewables 2021: Global status report. Technical report, Renewable Energy Policy Network for the 21st
751 Century, 1 rue Miollis, Building VII, 75015 Paris, France.
- 752 Rigollier, C., Lefevre, M., and Wald, L. (2004). The method Heliosat-2 for deriving shortwave solar radiation from satellite
753 images. *Solar Energy*, 77(2):159–169.
- 754 Rombach, R., Blattmann, A., Lorenz, D., Esser, P., and Ommer, B. (2022). High-resolution image synthesis with latent diffusion
755 models. In *2022 IEEE/CVF Conference on Computer Vision and Pattern Recognition (CVPR)*, pages 10674–10685.
- 756 Ronneberger, O., Fischer, P., and Brox, T. (2015). U-Net: Convolutional networks for biomedical image segmentation. In
757 Navab, N., Hornegger, J., Wells, W. M., and Frangi, A. F., editors, *Medical Image Computing and Computer-Assisted
758 Intervention – MICCAI 2015*, pages 234–241, Cham. Springer International Publishing.
- 759 Salio, P., Nicolini, M., and Zipser, E. J. (01 Apr. 2007). Mesoscale convective systems over southeastern south america and
760 their relationship with the south american low-level jet. *Monthly Weather Review*, 135(4):1290–1309.
- 761 Schneider, R., Bonavita, M., Geer, A., Arcucci, R., Dueben, P., Vitolo, C., Le Saux, B., Demir, B., and Mathieu, P.-P.
762 (2022). ESA-ECMWF report on recent progress and research directions in machine learning for Earth system observation
763 and prediction. *npj Climate and Atmospheric Science*, 5(1):1–5.
- 764 Seity, Y., Brousseau, P., Malardel, S., Hello, G., Bénard, P., Bouttier, F., Lac, C., and Masson, V. (2011). The AROME-France
765 convective-scale operational model. *Monthly Weather Review*, 139(3):976–991.
- 766 Smith, A. G., Petersen, J., Selvan, R., and Rasmussen, C. R. (2020). Segmentation of roots in soil with U-Net. *Plant Methods*,
767 16.
- 768 Su, X., Li, T., An, C., and Wang, G. (2020). Prediction of Short-Time Cloud Motion Using a Deep-Learning Model. *Atmosphere*,
769 11(11).
- 770 Sánchez, J., Meinhardt-Llopis, E., and Facciolo, G. (2013). TV-L1 optical flow estimation. *Image Processing Online (IPOL)*.
- 771 Sønderby, C. K., Espeholt, L., Heek, J., Dehghani, M., Oliver, A., Salimans, T., Agrawal, S., Hickey, J., and Kalchbrenner, N.
772 (2020). MetNet: A neural weather model for precipitation forecasting. *arXiv preprint arXiv:2003.12140*.
- 773 Tan, C., Feng, X., Long, J., and Geng, L. (2018). FORECAST-CLSTM: A new convolutional LSTM network for cloudage
774 nowcasting. In *2018 IEEE Visual Communications and Image Processing (VCIP)*, pages 1–4.
- 775 Urbich, I., Bendix, J., and Müller, R. (2019). The Seamless Solar Radiation (SESORA) Forecast for Solar Surface Irradi-
776 ance—Method and Validation. *Remote Sensing*, 11(21).
- 777 Uwe, P., Steffen, K., Trentmann, Jörg, Rainer, H., Petra, F., Johannes, K., and Martin, W. (2019). Surface radiation data set
778 - heliosat (sarah) - edition 2.1.
- 779 Voyant, C., Notton, G., Kalogirou, S., Nivet, M.-L., Paoli, C., Motte, F., and Fouilloy, A. (2017). Machine learning methods

780 for solar radiation forecasting: A review. *Renewable Energy*, 105:569–582.

781 Wang, Z., Bovik, A., Sheikh, H., and Simoncelli, E. (2004). Image quality assessment: from error visibility to structural
782 similarity. *IEEE Transactions on Image Processing*, 13(4):600–612.

783 Wei, S., Zhang, H., Wang, C., Wang, Y., and Xu, L. (2019). Multi-temporal sar data large-scale crop mapping based on u-net
784 model. *Remote Sensing*, 11(1).

785 Yang, D., Alessandrini, S., Antonanzas, J., Antonanzas-Torres, F., Badescu, V., Beyer, H. G., Blaga, R., Boland, J., Bright,
786 J. M., Coimbra, C. F., David, M., Âzeddine Frimane, Gueymard, C. A., Hong, T., Kay, M. J., Killinger, S., Kleissl, J.,
787 Lauret, P., Lorenz, E., van der Meer, D., Paulescu, M., Perez, R., Perpiñán-Lamigueiro, O., Peters, I. M., Reikard, G., Renné,
788 D., Saint-Drenan, Y.-M., Shuai, Y., Urraca, R., Verbois, H., Vignola, F., Voyant, C., and Zhang, J. (2020). Verification of
789 deterministic solar forecasts. *Solar Energy*, 210:20–37. Special Issue on Grid Integration.

790 Yang, D., Wang, W., Gueymard, C. A., Hong, T., Kleissl, J., Huang, J., Perez, M. J., Perez, R., Bright, J. M., Xia, X., van der
791 Meer, D., and Peters, I. M. (2022). A review of solar forecasting, its dependence on atmospheric sciences and implications
792 for grid integration: Towards carbon neutrality. *Renewable and Sustainable Energy Reviews*, 161:112348.

793 Zach, C., Pock, T., and Bischof, H. (2007). A duality based approach for realtime TV-L1 optical flow. In *Proceedings of the*
794 *29th DAGM Conference on Pattern Recognition*, pages 214–223, Berlin, Heidelberg. Springer-Verlag.

795 Zhao, H., Gallo, O., Frosio, I., and Kautz, J. (2017). Loss functions for image restoration with neural networks. *IEEE*
796 *Transactions on Computational Imaging*, 3(1):47–57.

797 Zhou, Z., Siddiquee, M., Tajbakhsh, N., and Liang, J. (2020). UNet++: Redesigning skip connections to exploit multiscale
798 features in image segmentation. *IEEE Transactions on Medical Imaging*, 39(6):1856—1867.